

**Sentinels**



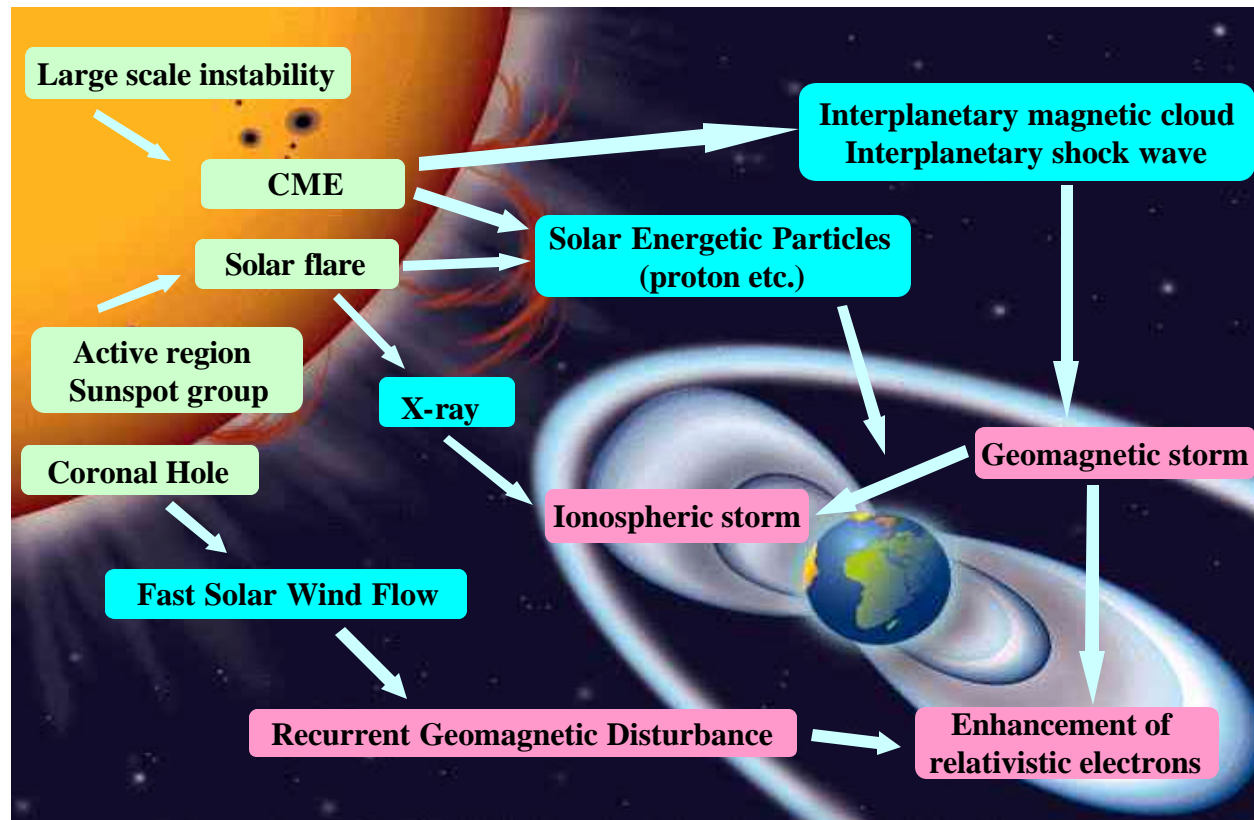
# **Living With a Star Sentinels**

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**LWS MOWG**

NASA HQ – June 19-20, 2003



Discover, understand and model the connection between solar phenomena and geospace disturbances.



- The LWS Science Architecture Team and Sentinels Pre-Formulation Team presented science objectives, observational strategies and possible mission scenarios to NASA HQ.

March 2000 – July 2001

- International LWS program initiated.

May 2001

- LWS SAT Findings and Report

July – September 2001

- First Heliospheric Strategy Panel Meeting.

November 2002

- HSP Report Completed

Fall 2003

- LWS Sentinels SDT formed

Early 2004



### Objectives:

Assist the LWS Sentinels Project Scientist by identifying the near-term strategy to make progress towards the LWS heliospheric science objectives. In particular:

- Identify heliospheric measurements required to accomplish the LWS program objectives.
- Determine to what extent current and future assets could be used to contribute toward the LWS heliospheric objectives and make recommendations for improved utilization.
- Identify current and past mission data sets relevant for the LWS heliospheric objectives that are not publicly available and suggest possible solutions.



## Primary

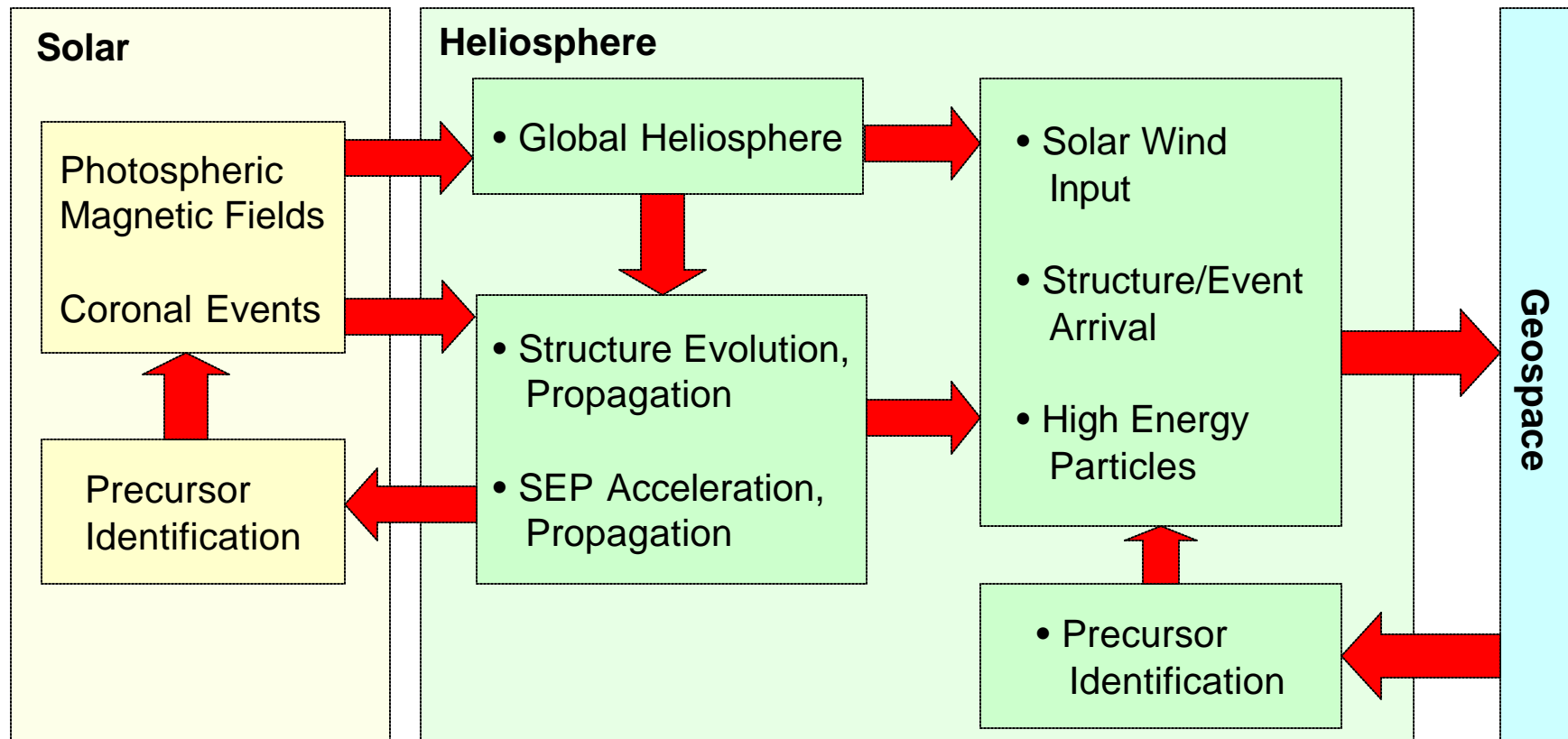
Stuart Bale	Berkeley
Nat Gopalswamy	GSFC
Dennis Haggerty	APL
Steve Kahler	AFRL
Davin Larson	Berkeley
Glenn Mason	U. Maryland
Chuck Smith	Bartol
John Steinberg	LANL
Adam Szabo	GSFC
Allan Tylka	NRL
Thomas Zurbuchen	U. Michigan

## Advisory

Joan Feynman	JPL
Len Fisk	U. Michigan
Tony Galvin	UNH
Bob Lin	Berkeley
David McComas	LANL
Dick Mewaldt	Caltech
Ed Roelof	APL

## Project

Nicky Fox	APL
Bob Hoffman	GSFC
Barbara Giles	GSFC
Vic Pizzo	NOAA
David Sibeck	GSFC
Chris StCyr	GSFC
Barbara Thompson	GSFC
Ron Zwickl	NOAA



- Identify gaps in current state of knowledge
- Determine areas of research where progress is most likely

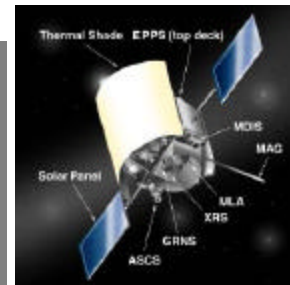
## Sentinels

## Preliminary Recommendations (1)

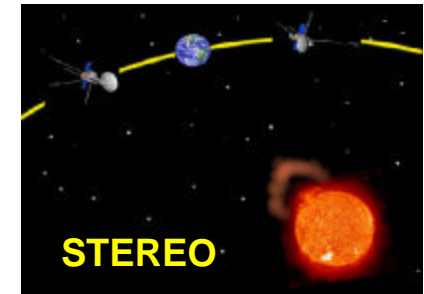


**When possible, fill observational gaps using future NASA missions:**

STEREO – Extend mission as late as possible  
MESSENGER – Collect cruise data



MESSENGER



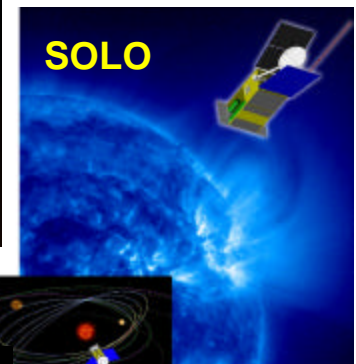
STEREO

**Cultivate international partnerships, ILWS:**

Solar Orbiter – *In situ* instrumentation, telemetry  
Bepi-Colombo – Interplanetary and cruise data  
Ulysses – Continuation of mission



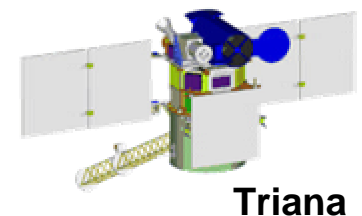
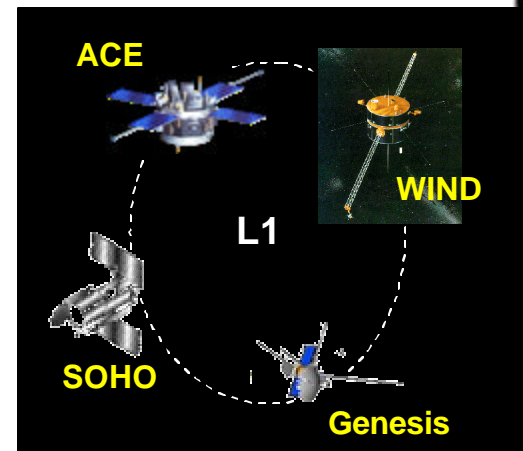
Bepi-Colombo



SOLO

**Extend the capabilities of existing near-Earth solar wind assets:**

L1 Cluster – Forge unified observatory from ACE, WIND, Genesis, SOHO  
Triana – Seek opportunities for launch



Triana



- **Establish SDT for US component to improve on current space holder:**

**Spacecraft:** 4 identical spinning s/c with at least five-year lifetime.

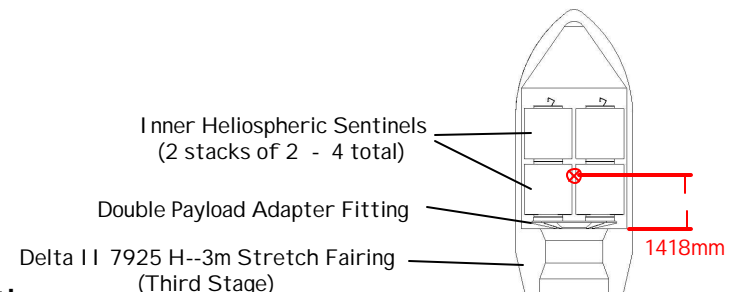
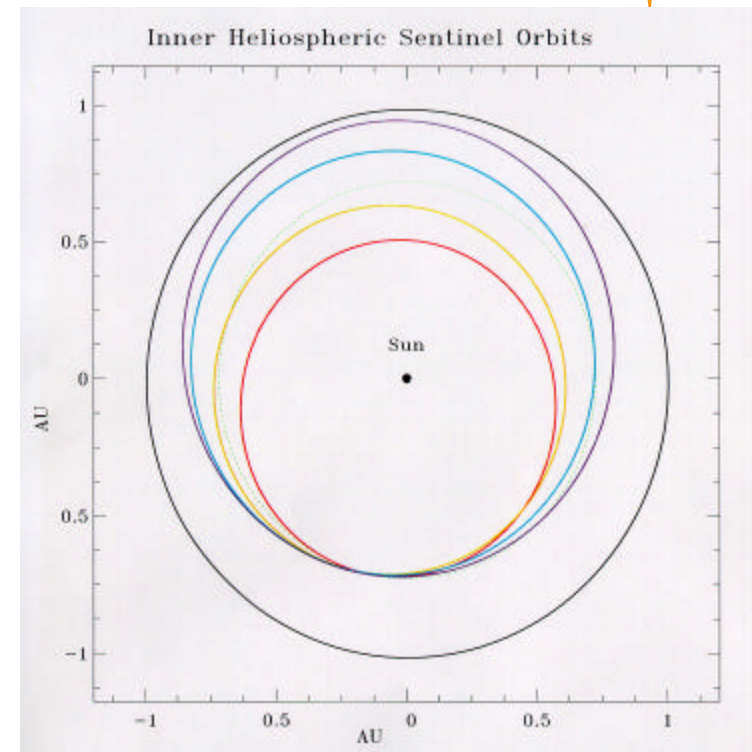
**Orbits:** 4 elliptical heliocentric orbits (0.5 – 0.95 x 0.72 AU) in the ecliptic.

**Instruments:** Magnetometer, solar wind analyzer, energetic particle detector, radio waves instrument.

**Space Access:** Single launch on a medium class ELV (Delta II 7925H)

**Cost:** \$300M - \$350M including launch vehicle and 2 years of operations.

**Limitations:** No coverage within 0.5 AU  
No latitudinal coverage outside of the ecliptic.

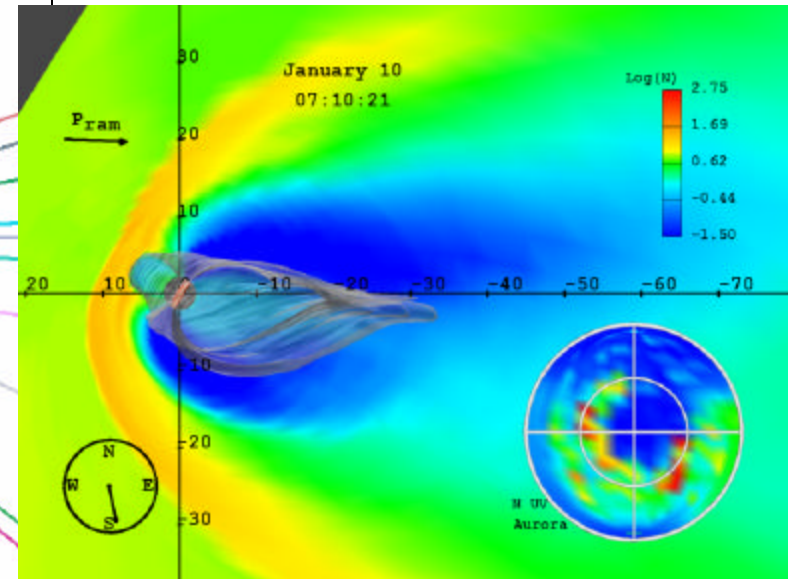
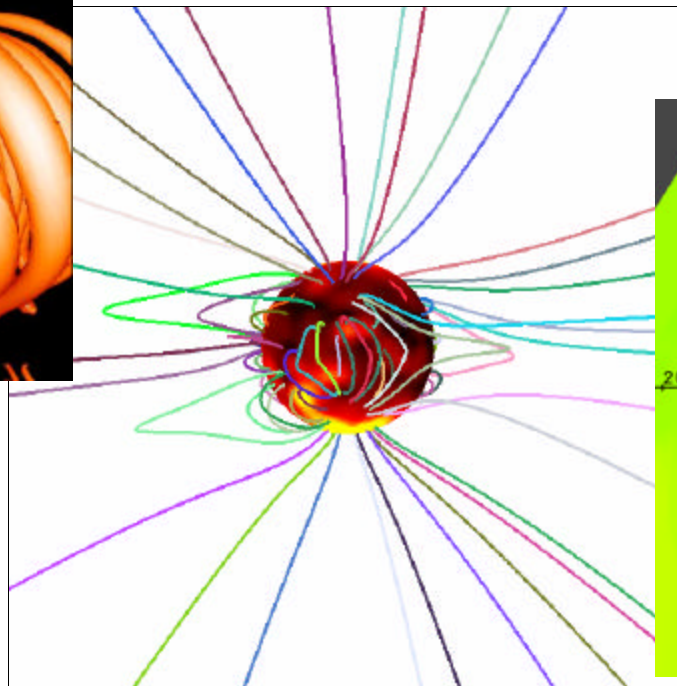
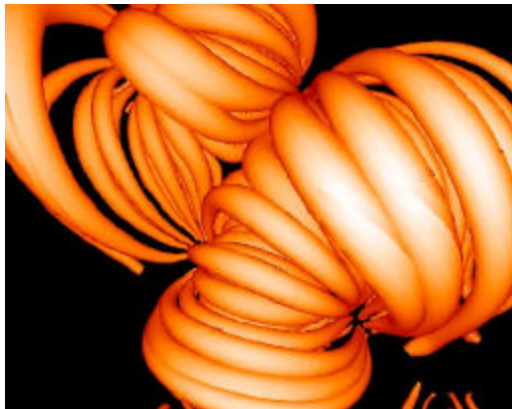






## The Role of Theory and Modeling

- Vigorous support of heliospheric modeling through TR&T particularly at the solar and magnetospheric interfaces
- Establish close ties between modeling center output and observational data services





## Data Services

- Develop a distributed data environment for all heliospheric missions, the Virtual Heliospheric Observatory (VHO)
- Treat observational data, processing service, and model output providers in similar fashion.
- Encourage interdisciplinary interaction.

